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On behalf of the PYTHIA8 collaboration

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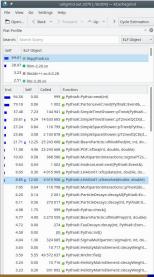
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Outline

- Introduction
- Recent developments
 - ► ParallelPythia, Shower variations, biasing
- New developments
 - Hadronisation weights, Onium showers
- Future developments
 - Sectorised merging histories, GPUs,



Is PYTHIA8 a bottleneck?

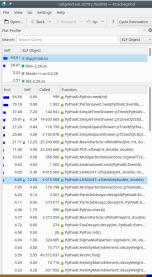


In its default setup PYTHIA8 is fast.

One (13 TeV $t\bar{t}$) event takes \sim 10 ms

- \sim 40% Final state shower
- \sim 30% Initial state shower
- ~ 10% Multi parton interactions
- ~ 10% Hadronisation + decays
- \sim 50k random numbers (\sim 6%)
- \sim 4k PDF evaluations (\sim 10%)

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- < 2% ME generation

Every step in PYTHIA8 depends on what happened in the previous step, so parallelisation is in general not possible.

Even where parallelisation is possible, the steps typically involves a lot of admin code that is not suited for e.g. GPUs.

But default PYTHIA8 is fast so no problem!



Paralellisation

Every event is independent, so PYTHIA8 itself can be parallelised

- Everything is perfectly thread safe
- OpenMP and HDF5 LHEF supported
- New wrapper class using std::thread



ParallelPythia

```
# from examples/main161.cc
#include "Pythia8/Pythia.h"
#include "Pythia8/PythiaParallel.h"
using namespace Pythia8:
int main() {
 // Use the PythiaParallel class for parallel generation.
 PvthiaParallel pvthia;
 pythia.readString("Beams:eCM = 8000.");
 pythia.readString("HardQCD:all = on");
 pythia.readString("PhaseSpace:pTHatMin = 20.");
 pvthia.init():
 Hist mult ("charged multiplicity", 100, -0.5, 799.5);
 // Use PythiaParallel::run to generate the specified number of events.
  pythia.run(10000, [&](Pythia* pythiaPtr) {
    // Find number of all final charged particles and fill histogram.
    int nCharged = 0:
    for (int i = 0; i < pvthiaPtr->event.size(); ++i)
     if (pvthiaPtr->event[i].isFinal() && pvthiaPtr->event[i].isCharged())
        ++nCharged;
    mult.fill( nCharged );
    // End of event loop. Statistics. Histogram. Done.
  }):
 pythia.stat();
  cout << mult:
  return 0:
```

Shower variations

With multiple weights per event we can save a lot of CPU time for scale variations etc.

Scale variations is easy for MEs. Showers are more tricky.

We have splitting functions and no emission functions

$$P(
ho_{\perp}^2) imes \exp\left(-\int_{
ho_{\perp}^2}^{
ho_{\perp, ext{prev}}^2} dk_{\perp}^2 P(k_{\perp}^2)
ight)$$

Using the veto algorithm we oversample with a simple overestimate function and throw to get the correct distribution.

Every time we accept/reject we can get weights for different variations,

$$P(p_{\perp}^2, lpha_{\mathcal{S}}(\mu_R))
ightarrow P(p_{\perp}^2, lpha_{\mathcal{S}}(C\mu_R)) \quad \Rightarrow \quad p_{ ext{acc}}
ightarrow p_{ ext{acc}}', p_{ ext{rej}}
ightarrow p_{ ext{rej}}',$$

accept/reject according p_{acc} and p_{rej} , and get an event weight

$$w_{ev} = \prod_i \frac{\rho_i'}{\rho_i}$$

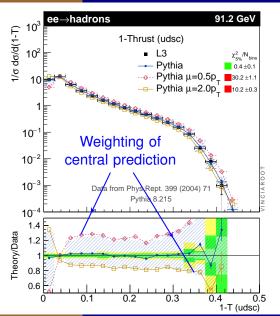
Note that large reweighting factors may result in fluctuating weights, so the statistical uncertainty will increase:

$$N_{ ext{ev, eff}} = rac{\left(\sum w
ight)^2}{\sum w^2}$$



- \blacktriangleright μ_B scale for QCD emissions in FSR
- \blacktriangleright μ_B scale for QCD emissions in ISR
- inclusion of non-singular terms in QCD emissions in FSR
- inclusion of non-singular terms in QCD emissions in ISR
- PDF members of a PDF family in LHAPDF6
- individual PDF members of a PDF family in LHAPDF6







Biasing

The same procedure may be used to bias your events, e.g. to get more $g \to b\bar{b}$

$$P'_{b\bar{b}}(p_\perp^2) = CP_{b\bar{b}}(p_\perp^2)$$

but now the rôle of variation and standard setting are reversed, so we accept/reject according to $p'_{\rm acc}$ and $p'_{\rm rej}$, and get an event weight

$$w_{ev} = \prod_i rac{
ho_i}{
ho_i'}$$



Biasing works well for low probability processes, but for more common processes we are hit by large weight fluctuations.

The statistical uncertainty of an observable scales like a power of the inverse no-emission probability, Δ^{-h}

For an *n*-emission observable we get an uncertainty

$$\delta \mathcal{O}_n \propto \frac{\Delta^{-h}}{\sqrt{C^n}}, \quad \text{with} \quad h \approx C$$

You cannot bias MPI.



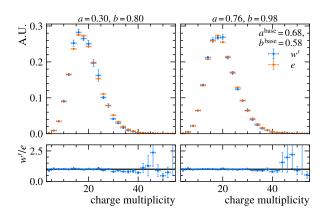
Hadronisation weights

In addition, weight variations are soon also available for hadronisation parameters.

Works on the same principle as shower variations (with the same caveats), but this time modifying hit-and-miss algorithms for e.g. the Lund symmetric fragmentation function.

$$p(z) \propto rac{(1-z)^a}{z} e^{-bm_{\perp}^2/z}$$

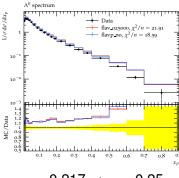




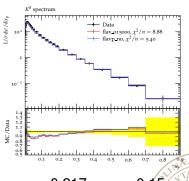


Flavour variations

 $\rho = StringFlav: probStoUD$

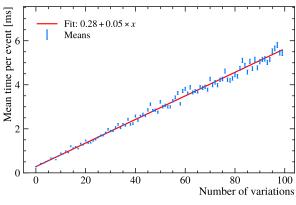


$$ho = 0.217
ightarrow
ho = 0.25$$



$$\rho = 0.217 \rightarrow \rho = 0.15$$

Timing





This is mainly useful for tuning

But again, if you can do weight variations you can do biasing (caveats still applies)



Onia showers

PYTHIA8 now includes shower splittings into charmonium and bottomonium splittings, e.g. $c \to J/\psi + c$ and $g \to O(^3P_J) + g$ (it's unfortunately a bit slow)

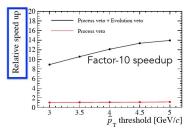
Biasing with a constant factor available.

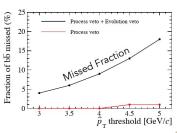


Biasing heavy flavours

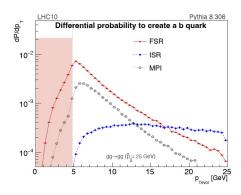
An alternative to biasing is to simply try to veto uninteresting events at a very early stage.

Using UserHooks you can tell PYTHIA8 to veto an event, e.g. when the evolution scale has dropped below some scale, \hat{p}_T .





Problem: aggressive speedup misses part of the b cross section



- ISR has the expected behaviour, FSR and MPI not.
- WIP: reformulate shower to have a more physical evolution scale so that all bs are produced at scales above m_b.

Beyond the default PYTHIA8

While the default PYTHIA8 is fast, we sometimes need speed-up.

In addition PYTHIA8 includes important optional modules which are quite CPU demanding, e.g.

- ► Hadronic rescattering has complexity N_h^2 (CPU×2)
- QCD-based Colour reconnections, N_d² or even N_d³ (CPU×15)
- ► CKKW-L type merging: N_I!



ME+PS merging

For CKKW-L style merging (UMEPS, NL3, UNLOPS, ...) you need to take all contributing shower histories into account for a given ME state.

In conventional PS each possible phase-space point receives contributions from many possible branching "histories" (or "clusterings"). Approximately sums over (singular) diagrams to give the full singularity structure.

# branchings	1	2	3	4	5	6	7
# histories	2	8	48	384	3840	46080	645120

histories \sim # feynman diagrams grow faster than factorial

Sector Showers

The default VINCIA shower in PYTHIA8 is unique in being a "Sector Shower"

- Divide N-gluon Phase Space into N "sectors", with step functions.
- Each PS sector corresponds to one specific gluon being the "softest" in the event — the one you would cluster if you were running a jet algorithm (specifically one called ARCLUS)
- Inside each sector, only a single kernel is allowed to contribute (the most singular one)!

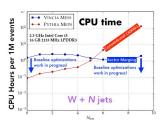
Sector Kernel = the eikonal for the soft gluon and its collinear DGLAP limits for z > 1/2.

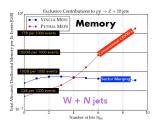
Only a single (product of) kernel(s) contributes to each phase-space point. $N! \rightarrow 1!$

[Skands & Villajero arXiv:1109.3608, Brooks, Preuss, Skands arXiv:2003.00702]

Sector Merging with VINCIA Sector Shower

Available since Pythia 8.306





Extentsons are coming:

- Sectorized matching at NNLO (proof of concept: arXiv:2108.07133, arXiv:2310.18671)
- Sectorized iterated tree-level ME corrections (demonstrated in arXiv:1109.3608)
- Sectorized multi-leg merging at NLO



Preview: VINCIA NNLO+PS for $H \rightarrow b\bar{b}$

Fixed-Order Reference = EERAD3 NLO $H \rightarrow b\bar{b}g$: already Highly optimised

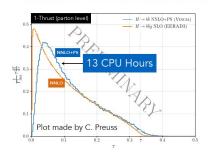


Uses analytical MEs, "folds" phase space to cancel azimuthally antipodal points, and uses antenna subtraction (\rightarrow smaller # of NLO subtraction terms than Catani-Seymour or FKS).

VINCIA NNLO + PS: uses the shower as phase-space generator: extremely efficient & everything formulated to be positive definite ⇒ no negative weights

Coloretti, Gehrmann-de Ridder, Preuss, JHEP 06 (2022) 009

➤ About factor 5 faster than EERAD3 (for comparable unweighted stats) + can be hadronised, etc.



Note:

NNLO accuracy in $H \to 2j$ implies NLO correction in first emission and LO correction in second emission.



Expectation: VINCIA NLO MEC approach order-of-magnitude faster than anything less optimised than EERAD3

GPU plans

Both hadronic rescattering and QCD-based colour reconnections can be improved algorithmically.

QCD-CR actively being worked on with M. Kreps (Warwick): Purely code-based modifications may result in speed increases of factors 2-3.

But some of the N^2 and N^3 calculations should be possible to parallelise and be delegated to a GPU.

This would in particular be important for heavy ion collisions, where $N \sim \mathcal{O}(1000)$ are common

Conclusions

PYTHIA8 may not be an important bottleneck, but there are improvements than can be made and are being made.



Other Thoughts (from Peter Skands)

Optimisation also crucial to reduce computational footprint / environmental impact

But funders do not (currently) score on this criterion at all

E.g., ERC allow a "Do No Significant Harm" statement — but assessors told (in boldface) to ignore it ARC does not even have such a statement. Not sure about other agencies...?

Tricky choice if one has to compromise on scientific ambition? Some thoughts on this in "Computational scientists should consider climate impacts and grant agencies should reward them", <u>P Skands, Nature Rev.Phys. 5 (2023) 3, 137-138</u>

All grants I am connected with now include minimisation of footprint as explicit goal

ARC DP22 "Tackling the computational bottleneck in precision particle physics" — on sector-based approaches

ARC DP23 "Beautiful Strings" — on more efficient (and better) models of heavy flavour production, fragmentation, and decays (incl matching), QED showers in hadron decays, collective effects in fragmentation, and Colour Reconnections:

POST DOC AT MONASH NOW OPEN FOR APPLICATIONS

Monash-Warwick Alliance for Particle Physics: including optimisation and improvements to EVTGEN and PYTHIA for HF physics (incl QED showers and QCD CR)

Royal Society Wolfson Visiting Fellowship "Piercing the precision barrier in high-energy particle physics": to develop efficient techniques for NNLO matching and beyond + interact with PanScales and with Warwick on (multithreaded) multipole QED showers in hadron decays

DECRA 23 [L. Scyboz, Monash]: "Bridging the accuracy gap: High-precision parton showers for colliders" — on PanScales and VINCIA